

*Dedicated to Professor Ioan-Iovitz Popescu's 80<sup>th</sup> Anniversary*

## GAMMA RESONANCES NEAR THRESHOLD FOR THE PRODUCTION OF THERMAL PHOTONEUTRONS

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*Received October 2, 2012*

*Abstract.* We have determined the positions of the  $(\gamma, n)$  resonances and upper limits for the integrated cross sections for the  $(\gamma, n)$  reactions, using data for the inverse process  $(n, \gamma)$ . We have estimated the number of low-energy neutrons which can be generated by the irradiation of a target with a  $\gamma$ -ray beam. Among the reactions producing thermal neutrons *via*  $(\gamma, n)$  reaction we mention  $^{185}\text{Re}(\gamma, n)^{184}\text{Re}$  with an upper limit of the integrated cross section of 2.4 b-eV.

*Key words:* gamma resonances, thermal photoneutrons, cross-section.

### 1. INTRODUCTION

In this work we consider an experiment in which a  $\gamma$ -ray beam induces nuclear electromagnetic transitions in a target, in the vicinity of the threshold  $S_n$  for the reaction  $(\gamma, n)$ .

We determine the positions of the  $(\gamma, n)$  resonances and upper limits for the integrated cross sections for the  $(\gamma, n)$  reactions, using data for the inverse process  $(n, \gamma)$ . We list the  $(\gamma, n)$  resonances and the upper limit for the  $(\gamma, n)$  reaction cross sections, in the vicinity of the reaction threshold. With the aid of these data we shall estimate the number of low-energy neutrons which can be generated by the irradiation of a target as a function of the spectral intensity of the  $\gamma$ -ray beam.

### 2. DETERMINATION OF $(\gamma, n)$ RESONANCES FROM THE INVERSE $(n, \gamma)$ PROCESS

The positions of the gamma-neutron resonances  $(\gamma, n)$  above threshold have been determined using existing data for the inverse process  $(n, \gamma)$ , with the aid of the relations between the variables characterizing the direct and inverse processes.

The positions of the ( $\gamma,n$ ) resonances and upper limits to the cross section of the ( $\gamma,n$ ) reaction have been determined for 73 stable isotopes and 101 unstable isotopes, for the energy intervals where we could find data for the ( $n,\gamma$ ) reactions.

We have used these results in order to identify the cases for which a ( $\gamma,n$ ) resonance has a position such that the neutron emitted backwards with respect to the direction of incidence of the gamma-ray photon may have zero energy [1].

Among the reactions for which there are such resonances we mention  $^{185}\text{Re}(\gamma,n)^{184}\text{Re}$  with an upper limit of the integrated cross section of 2.4 b-eV, and  $^{178}\text{Hf}(\gamma,n)^{177}\text{Hf}$  with an upper limit of the integrated cross section of 0.9 b-eV, in all directions.

## 2.1. STABLE ISOTOPES

The stable nuclei for which we could find thermal-neutron resonances are listed in Table 1 up to energies of the emitted neutrons of 0.1 eV. Upper limits for the corresponding integrated cross sections, with the emission of neutrons into  $4\pi$ , are also given in Table 1.

Table 1

Stable-isotope ( $\gamma,n$ ) resonances and upper limits for the integrated cross sections for the ( $\gamma,n$ ) reaction, with the emission of the neutrons into  $4\pi$ .  $D_-$  represents the energy of the neutron emitted backwards with respect to the direction of incidence, and  $D_+$  the energy of the neutron emitted along the direction of incidence of the  $\gamma$ -ray photon

Reaction	Upper limit for the integrated cross section, $B[-\text{eV}]$	Position of resonance $E_\gamma - S_n$ , [eV]	Width of resonance, [eV]	Height of resonance, $b$	Lower limit of the neutron energy $D_-$ , [eV]	Upper limit of the neutron energy $D_+$ , [eV]
$^{153}\text{Eu}(\gamma,n)^{152}\text{Eu}$	0.007	258.2	0.187	4.973E-2	1.588E-4	6.704
$^{170}\text{Yb}(\gamma,n)^{169}\text{Yb}$	0.036	227.9	0.145	1.666E-1	3.348E-4	5.295
$^{185}\text{Re}(\gamma,n)^{184}\text{Re}$	2.438	171.5	0.151	1.081E1	9.013E-4	3.837
$^{152}\text{Sm}(\gamma,n)^{151}\text{Sm}$	0.121	242.6	0.124	7.559E-1	9.863E-4	6.557
$^{178}\text{Hf}(\gamma,n)^{177}\text{Hf}$	0.971	176.5	0.088	7.657E0	0.00197	4.156
$^{153}\text{Eu}(\gamma,n)^{152}\text{Eu}$	0.044	258.4	0.224	1.462E-1	0.00396	7.101
$^{156}\text{Gd}(\gamma,n)^{155}\text{Gd}$	0.069	252.8	0.150	3.176E-1	0.01791	7.190
$^{153}\text{Eu}(\gamma,n)^{152}\text{Eu}$	0.010	257.9	0.281	3.756E-2	0.02136	6.030
$^{152}\text{Sm}(\gamma,n)^{151}\text{Sm}$	0.212	242.9	0.116	1.299E0	0.02259	7.181
$^{154}\text{Sm}(\gamma,n)^{153}\text{Sm}$	1.735	223.2	0.106	1.188E1	0.02957	6.659
$^{193}\text{Ir}(\gamma,n)^{192}\text{Ir}$	0.059	168.5	0.106	3.786E-1	0.04058	2.798
$^{164}\text{Dy}(\gamma,n)^{163}\text{Dy}$	0.653	193.7	0.126	3.501E0	0.04557	5.699
$^{152}\text{Sm}(\gamma,n)^{151}\text{Sm}$	0.270	242.0	0.102	1.811E0	0.05238	5.292
$^{139}\text{La}(\gamma,n)^{138}\text{La}$	0.091	300.7	0.142	4.502E-1	0.06052	10.155
$^{153}\text{Eu}(\gamma,n)^{152}\text{Eu}$	0.014	257.4	0.142	6.860E-2	0.06719	5.488
$^{170}\text{Yb}(\gamma,n)^{169}\text{Yb}$	0.210	227.4	0.098	1.307E0	0.07073	4.216

Table 1 (continued)

$^{193}\text{Ir}(\gamma,n)^{192}\text{Ir}$	0.102	169.5	0.114	6.265E-1	0.09640	4.773
$^{170}\text{Yb}(\gamma,n)^{169}\text{Yb}$	0.227	228.7	0.109	1.204E0	0.09671	6.919
$^{150}\text{Sm}(\gamma,n)^{149}\text{Sm}$	0.417	229.2	0.082	3.358E0	0.09758	4.694

The  $(\gamma,n)$  resonances for several representative nuclei are shown in Figs. 1–3.

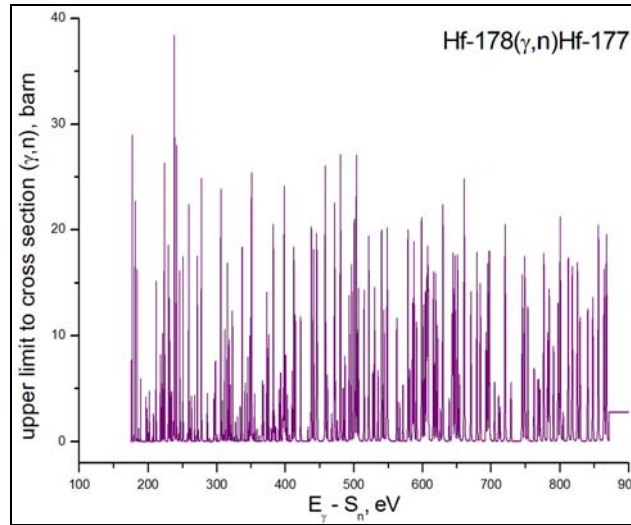


Fig. 1 –  $(\gamma,n)$  resonances and upper limits to the integrated cross section for the reaction  $^{178}\text{Hf}(\gamma,n)^{177}\text{Hf}$ .

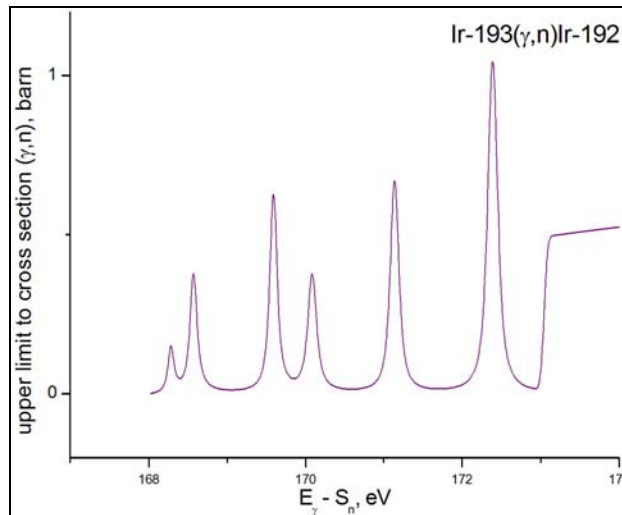


Fig. 2 –  $(\gamma,n)$  resonances and upper limits to the integrated cross section for the reaction  $^{193}\text{Ir}(\gamma,n)^{192}\text{Ir}$ .

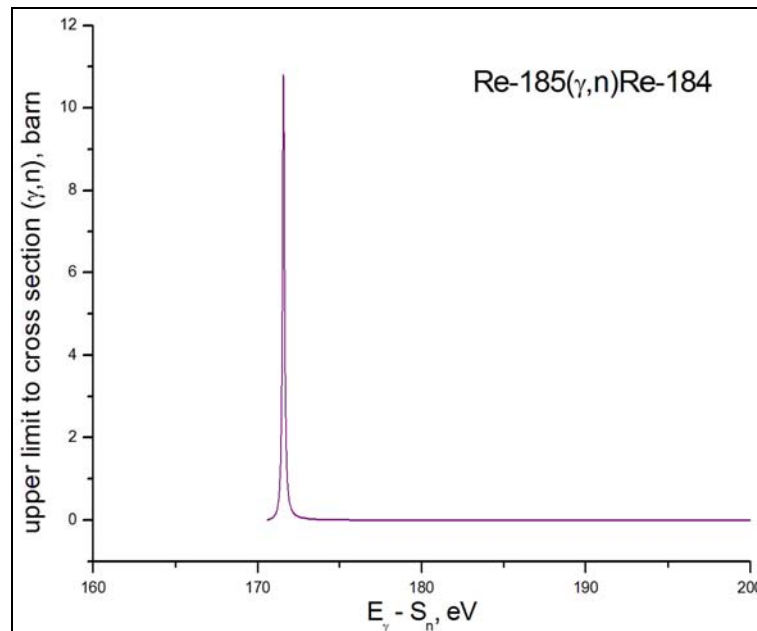


Fig. 3 –  $(\gamma,n)$  resonances and upper limits to the integrated cross section for the reaction  $^{185}\text{Re}(\gamma,n)^{184}\text{Re}$ .

## 2.2. UNSTABLE ISOTOPES

The unstable nuclei for which we could find thermal-neutron resonances are listed in Table 2 up to energies of the emitted neutrons of 0.1 eV. Upper limits for the corresponding integrated cross sections, with the emission of neutrons into  $4\pi$ , are also given in Table 2.

The upper limits for the integrated cross section correspond to the emission of the neutrons in all directions. The integrated cross sections for the generation of thermal neutrons are smaller than these limits by a factor of about 1/36, and can also be smaller than the listed limits because of the branching ratios.

## 3. GENERATION OF THERMAL NEUTRONS VIA $(\gamma,n)$ REACTIONS

The upper limit for the number of thermal neutrons generated per second *via* the  $(\gamma,n)$  reaction for a target having a thickness of  $3 \times 10^{-2}$  cm and for 400 incident gamma-ray photons/eV/s, a spectral intensity given in ref. [2], is of the order of  $(3 \times 10^{22} \text{ nuclei/cm}^3) \times (2 \text{ barn-eV}) \times (3 \times 10^{-2} \text{ cm}) \times (400 \text{ photons/eV/s}) = 0.72$  thermal neutrons/second, in all directions. The thermal neutrons represent a small fraction of the total number of generated photoneutrons.

Table 2

Unstable-isotope ( $\gamma,n$ ) resonances and upper limits for the integrated cross sections for the ( $\gamma,n$ ) reaction, with the emission of the neutrons into  $4\pi$ .  $D$  represents the energy of the neutron emitted backwards with respect to the direction of incidence, and  $D_+$  the energy of the neutron emitted along the direction of incidence of the  $\gamma$ -ray photon

Reaction	Upper limit for the integrated cross section, $B$ [-eV]	Position of resonance $E_\gamma - S_n$ , [eV]	Width of resonance [eV]	Height of resonance $b$	Lower limit of the neutron energy $D_-$ , [eV]	Upper limit of the neutron energy $D_+$ , [eV]	Half-life of target nuclei
$^{242}\text{Am}(\gamma,n)^{241}\text{Am}$	0.014	68.3	0.056	1.786E-1	4.317E-4	1.178	16.02 h
$^{243}\text{Am}(\gamma,n)^{242}\text{Am}$	7.370E-4	89.9	0.589	1.090E-3	4.887E-4	1.539	7370 y
$^{250}\text{Bk}(\gamma,n)^{249}\text{Bk}$	0.025	53.2	0.042	4.060E-1	4.915E-4	0.814	3.212 h
$^{192}\text{Ir}(\gamma,n)^{191}\text{Ir}$	0.025	108.1	0.086	1.946E-1	0.00272	2.417	73.827d
$^{152}\text{Eu}(\gamma,n)^{151}\text{Eu}$	0.021	141.6	0.110	1.241E-1	0.00306	3.948	13.537 y
$^{155}\text{Eu}(\gamma,n)^{154}\text{Eu}$	0.026	231.6	0.155	1.172E-1	0.00373	5.699	4.753 y
$^{233}\text{Pa}(\gamma,n)^{232}\text{Pa}$	0.002	98.5	0.369	5.990E-3	0.00449	1.529	26.975 d
$^{232}\text{Pa}(\gamma,n)^{231}\text{Pa}$	0.007	71.6	0.057	8.812E-2	0.00548	1.408	1.32 d
$^{240}\text{Pu}(\gamma,n)^{239}\text{Pu}$	0.010	95.8	0.100	5.723E-2	0.00778	1.388	6561 y
$^{230}\text{Th}(\gamma,n)^{229}\text{Th}$	0.060	108.3	0.060	7.029E-1	0.00802	2.143	75400 y
$^{242}\text{Pu}(\gamma,n)^{241}\text{Pu}$	0.010	88.5	0.114	5.689E-2	0.00885	1.252	375000 y
$^{236}\text{U}(\gamma,n)^{235}\text{U}$	0.002	97.7	0.240	4.610E-3	0.01164	1.398	2.342E7 y
$^{244}\text{Am}(\gamma,n)^{243}\text{Am}$	1.639E-4	63.8	0.095	1.100E-3	0.01745	1.334	10.1 h
$^{254}\text{Cf}(\gamma,n)^{253}\text{Cf}$	0.035	77.3	0.129	1.719E-1	0.01823	1.536	60.5 d
$^{232}\text{Pa}(\gamma,n)^{231}\text{Pa}$	0.002	71.7	0.060	2.572E-2	0.02044	1.577	1.32 d
$^{237}\text{Np}(\gamma,n)^{236}\text{Np}$	4.203E-4	98.2	0.271	1.560E-3	0.02151	1.310	2.144E6 y
$^{114}\text{In}(\gamma,n)^{113}\text{In}$	0.059	251.1	0.115	3.670E-1	0.02442	7.926	71.9s
$^{233}\text{Pa}(\gamma,n)^{232}\text{Pa}$	0.001	98.8	0.531	3.400E-3	0.02532	2.140	26.975 d
$^{199}\text{Au}(\gamma,n)^{198}\text{Au}$	1.113	155.7	0.202	3.597E0	0.02565	2.604	3.139 d
$^{238}\text{Np}(\gamma,n)^{237}\text{Np}$	0.006	68.4	0.052	8.015E-2	0.02597	1.522	2.117 d
$^{116}\text{In}(\gamma,n)^{115}\text{In}$	0.870	214.6	0.108	5.823E0	0.02717	6.549	14.1 s
$^{156}\text{Eu}(\gamma,n)^{155}\text{Eu}$	1.826	139.0	0.101	1.220E1	0.03027	2.951	15.19 d
$^{169}\text{Yb}(\gamma,n)^{168}\text{Yb}$	0.043	150.4	0.076	3.721E-1	0.03128	2.940	32.018 d
$^{237}\text{Np}(\gamma,n)^{236}\text{Np}$	0.004	98.7	0.141	1.634E-2	0.03578	2.191	2.144E6 y
$^{250}\text{Cf}(\gamma,n)^{249}\text{Cf}$	0.121	94.9	0.162	4.775E-1	0.04735	2.103	13.08 y
$^{147}\text{Pm}(\gamma,n)^{146}\text{Pm}$	1.775	215.3	0.148	7.950E0	0.04798	4.871	2.6234 y
$^{242}\text{Am}(\gamma,n)^{241}\text{Am}$	0.023	68.6	0.060	2.724E-1	0.04980	1.658	16.02 h
$^{227}\text{Ra}(\gamma,n)^{226}\text{Ra}$	0.002	49.7	0.045	3.234E-2	0.06853	1.432	42.2 m
$^{252}\text{Cf}(\gamma,n)^{251}\text{Cf}$	0.148	81.8	0.134	3.736E-1	0.07371	1.990	2.645 y
$^{152}\text{Eu}(\gamma,n)^{151}\text{Eu}$	0.058	141.0	0.101	4.454E-1	0.08511	2.689	13.537 y
$^{254}\text{Es}(\gamma,n)^{253}\text{Es}$	0.079	55.4	0.038	1.551E0	0.08733	1.508	275.7 d
$^{232}\text{Pa}(\gamma,n)^{231}\text{Pa}$	0.001	72.0	0.062	1.254E-2	0.09127	2.002	1.32 d
$^{155}\text{Eu}(\gamma,n)^{154}\text{Eu}$	0.004	231.1	0.164	1.792E-2	0.09151	4.605	4.753y
$^{177}\text{Lu}(\gamma,n)^{176}\text{Lu}$	0.074	153.3	0.089	6.036E-1	0.09840	4.725	6.647 d

#### 4. CONCLUSIONS

In this work we have determined the positions of the  $(\gamma,n)$  resonances and upper limits for the integrated cross sections for the  $(\gamma,n)$  reactions with the emission of neutrons in  $4\pi$ , using data for the inverse process  $(n,\gamma)$ .

We have identified the cases for which a  $(\gamma,n)$  resonance has a position such that the neutron emitted backwards with respect to the direction of incidence of the gamma-ray photon may have zero energy, and we have estimated the rate of generation of thermal neutrons *via* this mechanism.

Among the reactions producing thermal neutrons *via*  $(\gamma,n)$  reaction we mention  $^{185}\text{Re}(\gamma,n)^{184}\text{Re}$  with an upper limit of the integrated cross section of 2.4 b-eV, and  $^{178}\text{Hf}(\gamma,n)^{177}\text{Hf}$  with an upper limit of the integrated cross section of 0.9 b-eV, in all directions.

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